The quest for θ_{13}

- A theoretical perspective -

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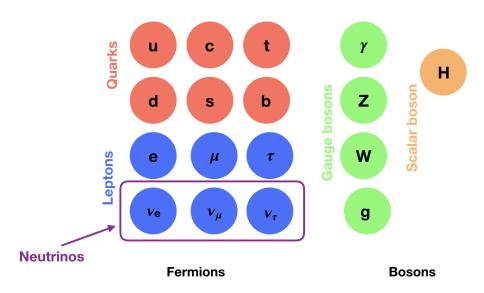




BNL Daya Bay Colloquium

15th December 2020

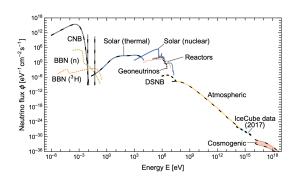
Neutrinos in the SM



Role of neutrinos in the Universe

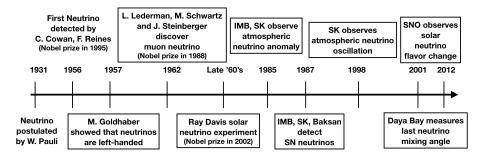
- second most abundant particle in the Universe:a billion times more neutrinos than matter particles in the Universe
- ► 10^{38} neutrinos per second produced by the Sun (flux of $\sim 10^{11}$ /cm²/sec at the Earth)
- 15 billion neutrinos from the Big Bang are in a large lecture hall





- neutrino energies span 24 orders of magnitude
- different neutrino sources and production mechanisms
- lots of room for exciting neutrino physics & lots to learn!

Neutrino discoveries & BNL's contribution



BNL had and has big impact on neutrino physics!

Neutrino oscillations

latest neutrino Nobel prize awarded to observation of flavor change of neutrinos:

only possible for massive neutrinos however neutrinos are massless in the SM

⇒ conclusive evidence for physics **beyond** the Standard Model!



Neutrino oscillations

flavor eigenstates (of weak interaction) and mass eigenstates (of free particle Hamiltonian) not aligned for neutrinos

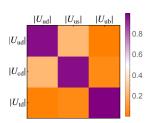
$$\begin{pmatrix} \nu_{e} \\ \nu_{\mu} \\ \nu_{\tau} \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} \begin{pmatrix} \nu_{1} \\ \nu_{2} \\ \nu_{3} \end{pmatrix}$$

U_{PMNS}: relates flavor and mass states

parametrized by 4 parameters (3 angles, 1 phase)

- Flavor mixing already present in the SM in the quark sector via the CKM matrix
- CKM matrix measured to a very good accuracy
 - \rightarrow small mixing angles in the quark sector, CKM matrix is close to a diagonal matrix

$$egin{aligned} & heta_{12}^{\textit{CKM}} = 13.1^{\circ}, \\ & heta_{13}^{\textit{CKM}} = 0.2^{\circ}, \\ & heta_{23}^{\textit{CKM}} = 2.3^{\circ} \end{aligned}$$



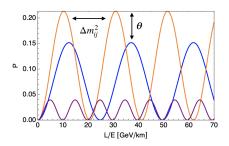
Shaped expectations for leptonic mixing matrix

Neutrino oscillations



produce neutrino of flavor α with energy E probability to detect neutrino with flavor β at distance L is

$$P(\nu_{\alpha} \rightarrow \nu_{\beta}) = \sin^2 2\theta \sin^2(\Delta m_{ij}^2 L/4E), \ \Delta m_{ij}^2 = m_i^2 - m_j^2$$



Neutrino oscillations

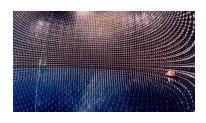
observation of neutrino oscillations introduced at least 6 new parameters to the Standard Model:

3 mixing angles, one CP violating phase, 2 mass splittings

⇒ **need** to measure them



Atmospheric mixing [hep-ex/9807003]



SuperKamiokande measured atmospheric mixing parameters:

$$\theta_{23} \approx 45^{\circ}$$

$$\Delta \textit{m}_{32}^2 \approx 2.2 \cdot 10^{-3} \; \text{eV}^2$$

large flavor mixing between ν_2 and ν_3

surprising if compared to small quark mixing!

Solar mixing [nucl-ex/0309004]



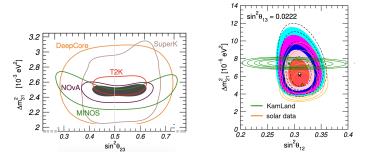
SNO measured solar mixing parameters:

$$\theta_{12}\approx 30.2^\circ$$

$$\Delta \textit{m}_{21}^2 \approx 5 \cdot 10^{-5} \; \text{eV}^2$$

large mixing between ν_1 , ν_2 !

SK & SNO measurements have been confirmed by other experiments like NOvA, MINOS, T2K, KamLand using man made neutrino sources like accelerator neutrinos or reactor neutrinos



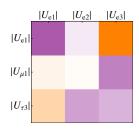
BNL contributed to these experiments (SK, SNO, MINOS)!

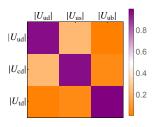
The last unknown mixing angle: θ_{13}

as of 2010: last unknown mixing angle θ_{13}

earlier hints and theoretical prejudice that this mixing angle is small!

→ lepton flavor models try to explain lepton mixing





Lepton flavor models

striking difference to quark mixing pattern: different origin of flavor

Quark mixing: small mixing angles, masses are hierarchical → continuous flavor symmetry

$$U_{CKM} \sim \mathbb{1} + \mathcal{O}(\lambda)$$

expansion parameter $\lambda \sim$ 0.22 (Wolfenstein parametrization)

[Wolfenstein '83]

Lepton mixing: large mixing angles

$$U_{PMNS} \sim U_0 \not\sim 1$$

non-diagonal U_0 can be realized with discrete flavor symmetries

discrete flavor symmetry as explanation for leptonic mixing example: tri-bimaximal mixing

$$U_{PMNS} = U_{TBM} = egin{pmatrix} \sqrt{2/3} & 1/\sqrt{3} & \textbf{0} \ -1/\sqrt{6} & 1/\sqrt{3} & 1/\sqrt{2} \ -1/\sqrt{6} & -1/\sqrt{3} & 1/\sqrt{2} \end{pmatrix}$$

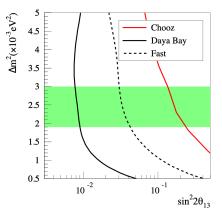
$$\theta_{23} = 45^{\circ}, \; \theta_{12} = \arcsin(1/\sqrt{3}) \approx 31.7^{\circ}, \, \theta_{13} = 0$$

many more models exist, they only differ by the predicted value of θ_{12}

 \Rightarrow theoretical expectation for θ_{13} shaped experimental program

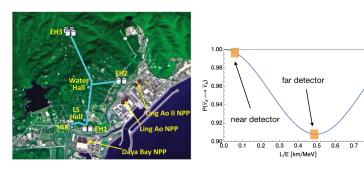
The last unknown mixing angle: θ_{13}

sensitivity of Daya Bay [hep-ex/0701029]



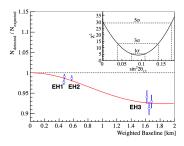
ightarrow Daya Bay had very good sensitivity to small angles $\sin^2 2\theta_{13} \sim \mathcal{O}(0.01) \approx 3^\circ$

Daya Bay experiment measures reactor anti electron neutrino survival probability



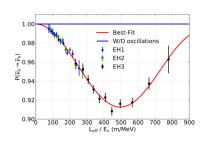
more details on the history of Daya Bay, BNL's involvement and the measurement will be covered in the following talks

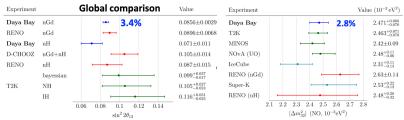
- 8 March 2012: Daya Bay publication using data from 24 December 2011-17 February 2012
- 6% deficit of electron neutrinos at far detector
 - \Rightarrow > 5 σ evidence for non-zero θ_{13}



publication in PRL has > 2000 citations!

Daya Bay results [1809.02261, Ling @Neutrino2020]





Daya Bay has world best precision on θ_{13}

Implications of θ_{13} measurement

 completed determination of leptonic mixing angles: reduced unknown number of SM parameters

confirmation of 3 flavor paradigm: agreement between atmospheric (ν_{μ}) and reactor (ν_{e}) mass splitting

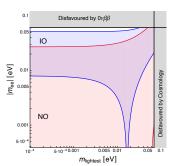
▶ determined fraction of ν_3 in electron neutrinos: reduce uncertainties in predicting neutrino related phenomena

Implications of θ_{13} : allowed parameter space for $|m_{ee}|$

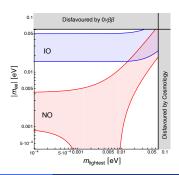
observable in neutrinoless double beta decay \Rightarrow determines if neutrinos are Majorana or Dirac particles

$$|m_{ee}| = \left| m_1 U_{e1}^2 + m_2 U_{e2}^2 + m_3 U_{e3}^2 \right|$$

before Daya Bay measurement: θ_{13} unknown



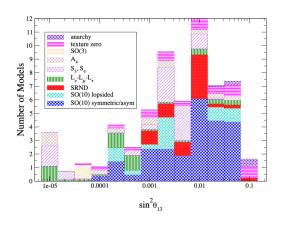
after Daya Bay measurement: θ_{13} known

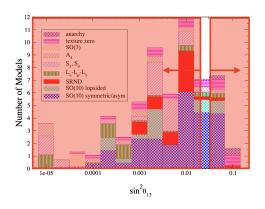


Implications of θ_{13} measurement

- completed determination of leptonic mixing angles: reduced unknown number of SM parameters
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- ightharpoonup determined fraction of ν_3 in electron neutrinos: reduce uncertainties in predicting neutrino related phenomena
- guide flavor theory model building

predictions of models to explain observed mixing pattern





many flavor models ruled out \rightarrow new ideas needed!

Julia Gehrlein (BNL) The quest for θ_{13} BNL Daya Bay Colloquium

Implications of θ_{13} measurement

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- guide flavor theory model building
- gateway to CP violation in lepton sector

Implications of θ_{13} : gateway to CPV

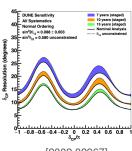
only with 3 non-zero mixing angles CP violation possible

$$P(\nu_{\alpha} \to \nu_{\beta}) - P(\overline{\nu}_{\alpha} \to \overline{\nu}_{\beta})$$

$$\propto \sin \theta_{12} \cos \theta_{12} \sin \theta_{23} \cos \theta_{23} \sin \theta_{13} \cos^{2} \theta_{13} \sin \delta$$

non-zero $\theta_{13} \Rightarrow$ possibility for CP violation in lepton sector!

next generation neutrino experiments like DUNE and T2HK will measure CPV and phase δ



[2002.02967]

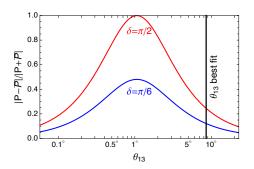
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- gateway to CP violation in lepton sector
- large θ_{13} : consequences for sensitivity of next generation of neutrino facilities to measure unknown neutrino parameters (δ and MO sensitivity)

Implications of θ_{13} : next generation neutrino experiments

Implications for the determination of unknown parameters

measurement of CP asymmetry for $P(\nu_{\mu} \rightarrow \nu_{e})$

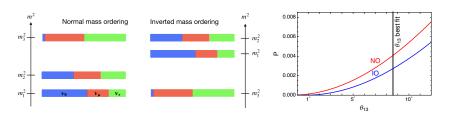


reasonable distinction power for different values of δ

Implications of θ_{13} : next generation neutrino experiments

Implications for the determination of unknown parameters

determination of mass ordering for $P(\nu_{\mu} \rightarrow \nu_{e})$



distinction between NO and IO increases with θ_{13}

Future of neutrino physics

- many open questions in neutrino physics: CPV, MO, neutrino mass scale, nature & origin of neutrino masses, new neutrino interactions, additional neutrino species...
- BNL well positioned for the future of neutrino physics:
 - experimental contributions to DUNE, nEXO, PROSPECT, SBN, Rubin Observatory LSST
 - theory developments: oscillation theory, model building, theoretical calculations for neutrino processes, ...

Thank you for your attention!

Backup: Neutrino observables

non-oscillation neutrino observables that depend on θ_{13} : observable in neutrinoless double beta decay:

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⇒ determine if neutrinos are Majorana or Dirac particles

Backup: Neutrino observables

non-oscillation neutrino observables that depend on θ_{13} : kinematic neutrino mass:

$$\begin{split} m_{\beta}^2 &= \sum_i |U_{ei}|^2 m_i^2 \\ &= m_1^2 \cos^2 \theta_{12} \cos^2 \theta_{13} + m_2^2 \sin^2 \theta_{12} \cos^2 \theta_{13} + m_3^2 \sin^2 \theta_{13} \end{split}$$

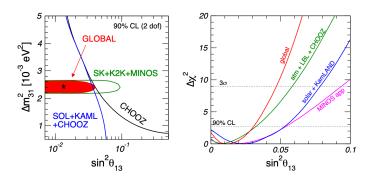
⇒ information on absolute neutrino mass scale

current bound $m_{\beta} < 1.1 \text{ eV}$

[KATRIN experiment '20]

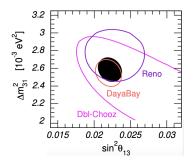
Backup: The last unknown mixing angle: θ_{13}

status May 2010 [1003.5800]



 \rightarrow only upper bounds on θ_{13} from atmospheric neutrino, reactor neutrino and long baseline experiments

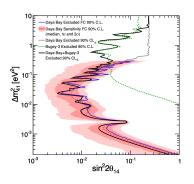
Backup: Daya Bay results [nuFit v5.0]



Daya Bay still has world best precision on θ_{13} & dominates global fits

Backup: Confirmation of 3 flavor paradigm

limits on sterile neutrinos



[2002.00301]

- > search for an additional oscillation frequency besides Δm_{32}^2 and Δm_{31}^2
- > set the most stringent upper limit for light sterile neutrino signal with $\Delta m_{A1}^2 < 0.2 \text{ eV}^2$

Backup: Implications of θ_{13} : flavor theory

many flavour models ruled out \rightarrow new ideas needed!

way out: introduce corrections to predictions from discrete symmetries $U_{PMNS} \sim U_0 + \mathcal{O}(\epsilon)$

interesting relation:

$$heta_{13}^{PMNS}pprox heta_{12}^{CKM}\sin heta_{23}^{PMNS}pprox 9.3^\circ$$

(can be realized if charged lepton mixing matrix has CKM like structure since $U_{PMNS}=U_l^\dagger U_\nu)$

precision measurement of all PMNS parameters key to test and disentangle discrete flavor symmetries!